

# Dimension Splitting Errors and a Long Time-Step Multi-Dimensional Scheme for Atmospheric Transport

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Meteorology



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The University of Reading

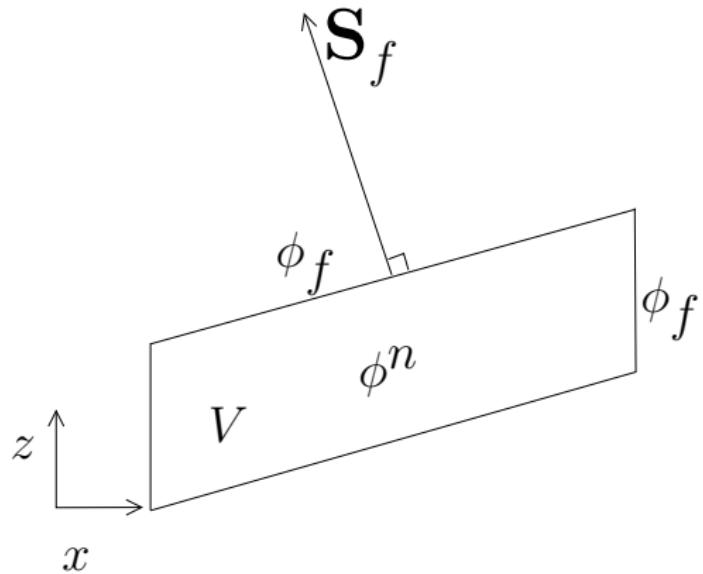
October 2016

# Motivation

- ▶ With cubed-sphere grid, Met-Office no longer need semi-Lagrangian for large Courant number over the poles
- ▶ Mass conservation
- ▶ Still need Courant number  $> 1$  in vertical
- ▶ Options for transport
  - ▶ Method of lines versus flux-form semi-Lagrangian (Forward in time)
  - ▶ Implicit, Explicit or HEVI
  - ▶ Dimensionally split or multi-dimensional

# Finite Volume Advection

$$\frac{\partial \phi}{\partial t} = -\nabla \cdot (\mathbf{u}\phi) - \frac{1}{V} \sum_{f \in \text{faces}} \phi_f \mathbf{u}_f \cdot \mathbf{S}_f$$

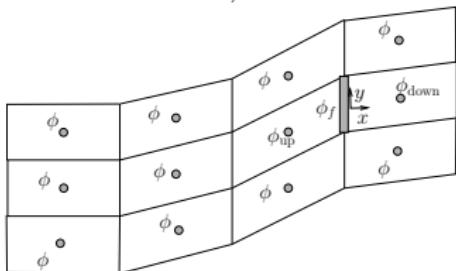


How to evaluate  $\phi_f$  between times  $n$  and  $n + 1$

# Method of Lines - Spatial Discretisation

- ▶ Method of lines - separate discretisation for space and time
  - ▶ At every instant, interpolate from surrounding cell values of  $\phi$  onto the face using an upwind-biased stencil:

27 cells in 3D, 12 in 2D:



Fit the polynomial:

$$a + bx + cx^2 + dx^3 +$$

$$ey + fxy + gx^2y$$

Second order assumptions:

- ▶ Cell averages = cell centre values
- ▶ Face averages = face centre values

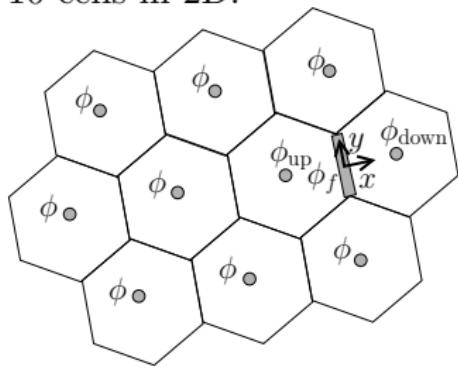
Use a least squares fit to find weights  $w_c$  for every cell in the stencil:

$$\phi_f = \phi_{up} + \sum_{c \in \text{stencil}} w_c \phi_c$$

Weights remain fixed every time-step

# Method of Lines - Spatial Discretisation

10 cells in 2D:



Fit the polynomial:

$$a + bx + cx^2 + dx^3 + \\ ey + fxy + gx^2y$$

# Method of Lines - can use Runge-Kutta Explicit Time Discretisation

Heun (2-stage, 2nd order RK):

$$\begin{aligned}\phi' &= \phi^n - \frac{\Delta t}{V} \sum_{f \in c} \left( \phi_{\text{up}}^n + \sum_{c \in \text{stencil}} w_c \phi_c^n \right) \mathbf{u}_f \cdot \mathbf{S}_f \\ \phi^{n+1} &= \frac{1}{2} \phi^n + \frac{1}{2} \phi' - \frac{\Delta t}{2V} \sum_{f \in c} \left( \phi'_{\text{up}} + \sum_{c \in \text{stencil}} w_c \phi'_c \right) \mathbf{u}_f \cdot \mathbf{S}_f\end{aligned}$$

## Method of Lines - Implicit Time Discretisation

Crank-Nicholson (as explicit correction of 1st-order upwind):

$$\phi' = \phi^n - \frac{\Delta t}{2V} \sum_{f \in c} \left( \phi_{\text{up}}^n + \phi'_{\text{up}} + 2 \sum_{c \in \text{stencil}} w_c \phi_c^n \right) \mathbf{u}_f \cdot \mathbf{S}_f$$

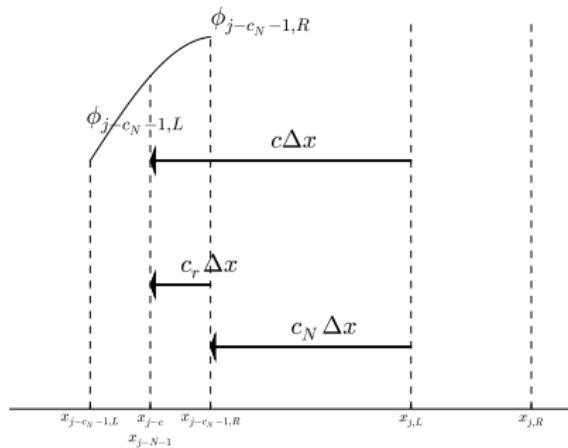
$$\phi^{n+1} = \phi^n - \frac{\Delta t}{2V} \sum_{f \in c} \left( \phi_{\text{up}}^n + \sum_{c \in \text{stencil}} w_c \phi_c^n + \phi_{\text{up}}^{n+1} + \sum_{c \in \text{stencil}} w_c \phi'_c \right) \mathbf{u}_f \cdot \mathbf{S}_f$$

## Method of Lines - HEVI

If the faces of cell  $c$  are separated into horizontal and vertical faces, this can be made HEVI:

$$\begin{aligned}\phi' &= \phi^n - \frac{\Delta t}{2V} \sum_{f \text{ vertical}} \left( \phi_{\text{up}}^n + \phi'_{\text{up}} + 2 \sum_{c \in \text{stencil}} w_c \phi_c^n \right) \mathbf{u}_f \cdot \mathbf{S}_f \\ &\quad - \frac{\Delta t}{V} \sum_{f \text{ horizontal}} \left( \phi_{\text{up}}^n + \sum_{c \in \text{stencil}} w_c \phi_c^n \right) \mathbf{u}_f \cdot \mathbf{S}_f \\ \phi^{n+1} &= \phi^n - \frac{\Delta t}{2V} \sum_{f \text{ vertical}} \left( \phi_{\text{up}}^n + \sum_{c \in \text{stencil}} w_c \phi_c^n + \phi_{\text{up}}^{n+1} + \sum_{c \in \text{stencil}} w_c \phi'_c \right) \mathbf{u}_f \cdot \mathbf{S}_f \\ &\quad - \frac{\Delta t}{2V} \sum_{f \text{ horizontal}} \left( \phi_{\text{up}}^n + \sum_{c \in \text{stencil}} w_c \phi_c^n + \phi'_{\text{up}} + \sum_{c \in \text{stencil}} w_c \phi'_c \right) \mathbf{u}_f \cdot \mathbf{S}_f\end{aligned}$$

# Dimensionally split Flux form semi-Lagrangian



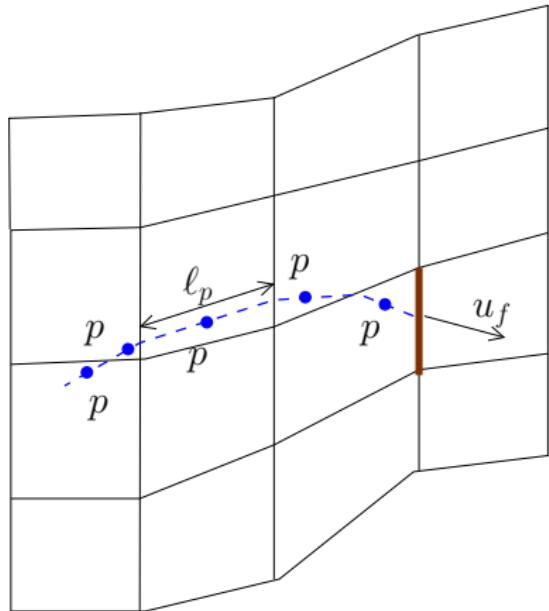
- ▶ COSMIC splitting (Leonard et al., 1996)

$$\phi_{ij}^{(n+1)} = \phi_{ij}^n + X_C \left( \phi_{ij}^{(n)} + \frac{1}{2} Y_A \left( \phi_{ij}^{(n)} \right) \right) + Y_C \left( \phi_{ij}^{(n)} + \frac{1}{2} X_A \left( \phi_{ij}^{(n)} \right) \right)$$

- ▶ PPM (Piecewise parabolic method Colella and Woodward, 1984) in each direction separately
- ▶ No monotonicity constraints
- ▶ Long time-steps
  - ▶ sum contributions from whole cells between face and departure point

# Multi-dimensional Flux form semi-Lagrangian

Integrate along a single trajectory for each face to find  $\phi_f$ , the average tracer swept through face  $f$



$$\phi_f = \sum_{p \in t} \phi_p \ell_p$$

Where  $\ell_p$  is the length of the part of the trajectory associated with point  $p$

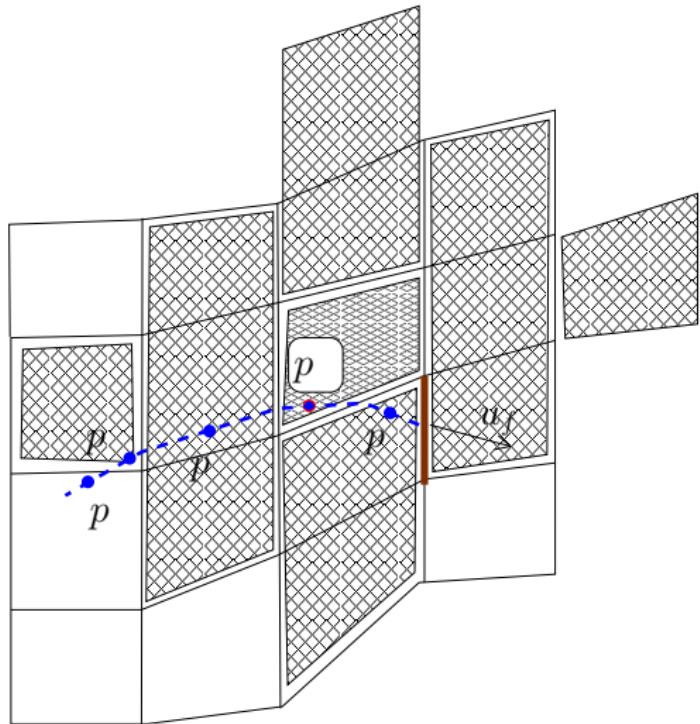
- ▶ Interpolate  $\phi$  onto all of the points  $p$  using a cubic polynomial with terms:

$$a + bx + cx^2 + dx^3 +  
ey + fxy + gx^2y +  
hy^2 + ixy^2 + jy^3$$

- ▶ The points  $p$  are mid-way along the part of the trajectory in each cell

# Multi-dimensional Flux form semi-Lagrangian

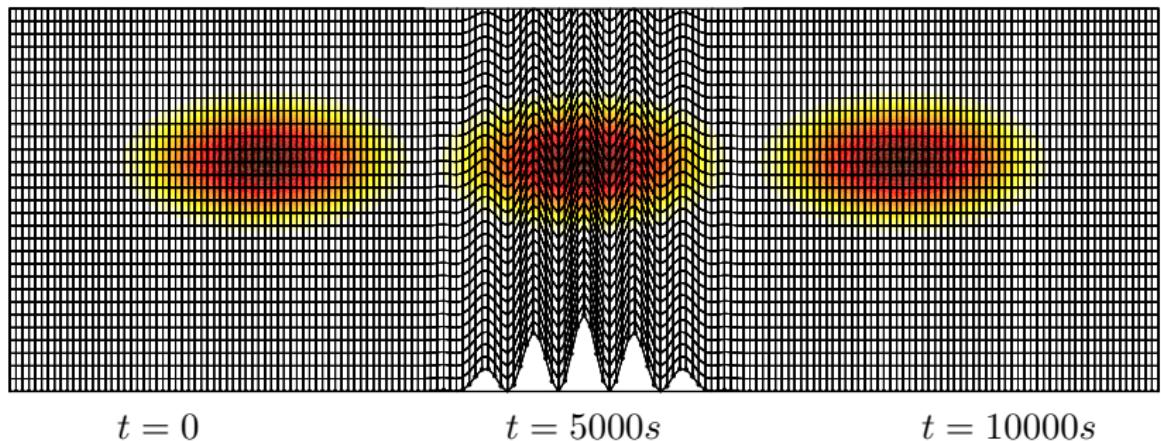
Stencil for interpolating onto each trajectory point,  $p$ ;



- ▶ Trajectory points, stencils and interpolation weights re-calculated every time the wind changes
- ▶ Very expensive

# Test Case: Horizontal Advection Over Orography

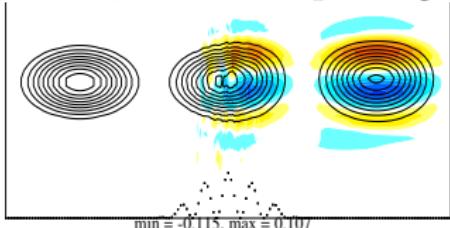
Schär et al, MWR, 2002:



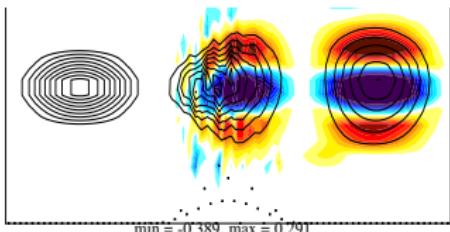
$\Delta x = 1\text{km}$ ,  $\Delta z = 500\text{m}$ ,  $u = 10\text{m/s}$ , mountain height,  $h_m = 3\text{km}$ ,  
tracer,  $50\text{km} \times 6\text{km}$

PPM, COSMIC splitting

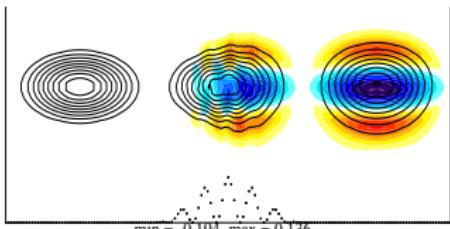
$\Delta t = 25$  s  
max  $c$ , 0.74  
mode  $c$ , 0.2  
 $c_d = 0.22$



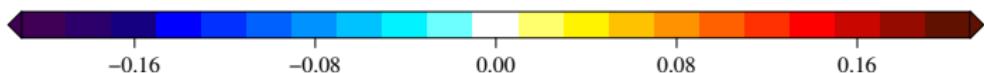
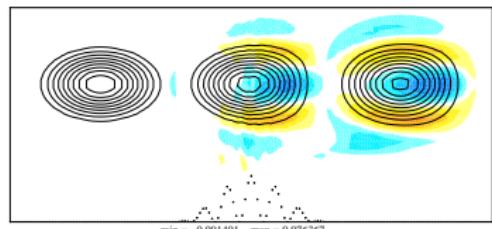
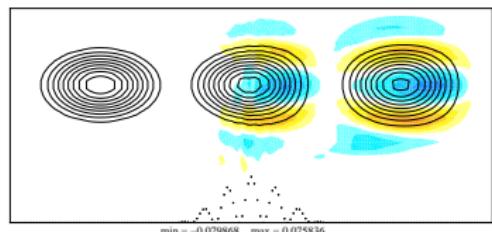
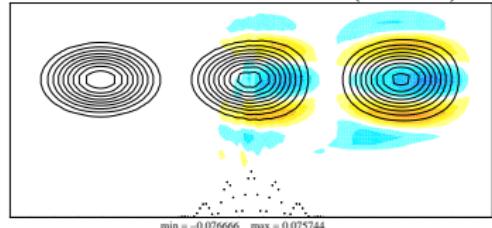
$\Delta t = 50$  s  
max  $c$ , 1.48  
mode  $c$ , 0.4  
 $c_d = 0.44$



$\Delta t = 100$  s  
max  $c$ , 2.96  
mode  $c$ , 0.8  
 $c_d = 0.87$



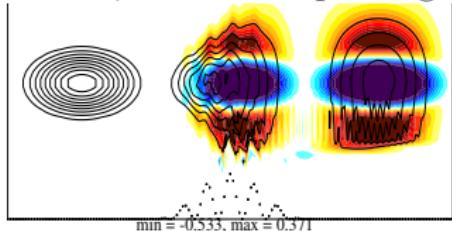
Multi-dimensional (MOL)



Contours from 0.1 to 0.9 every 0.1. Errors shaded.

PPM, COSMIC splitting

$\Delta t = 200$  s  
max  $c$ , 5.93  
mode  $c$ , 1.6  
 $c_d = 1.76$

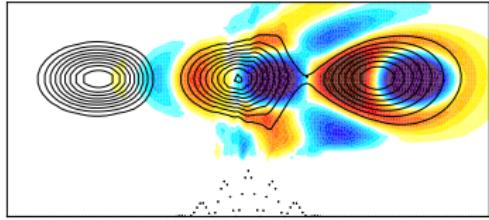


Multi-dimensional (MOL)

min = -0.15326 max = 0.083171

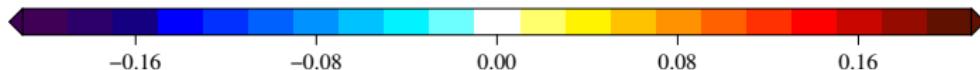
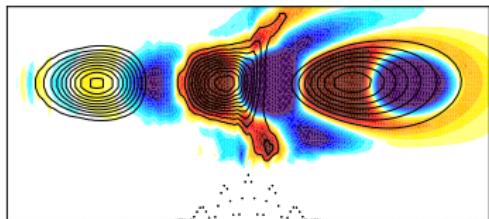
$\Delta t =$   
500 s  
max  $c$ , 15  
mode  $c$ , 4  
 $c_d = 4.4$

unstable



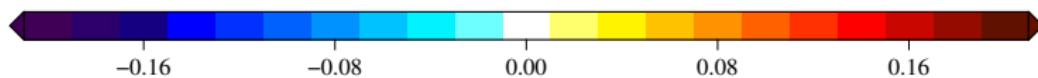
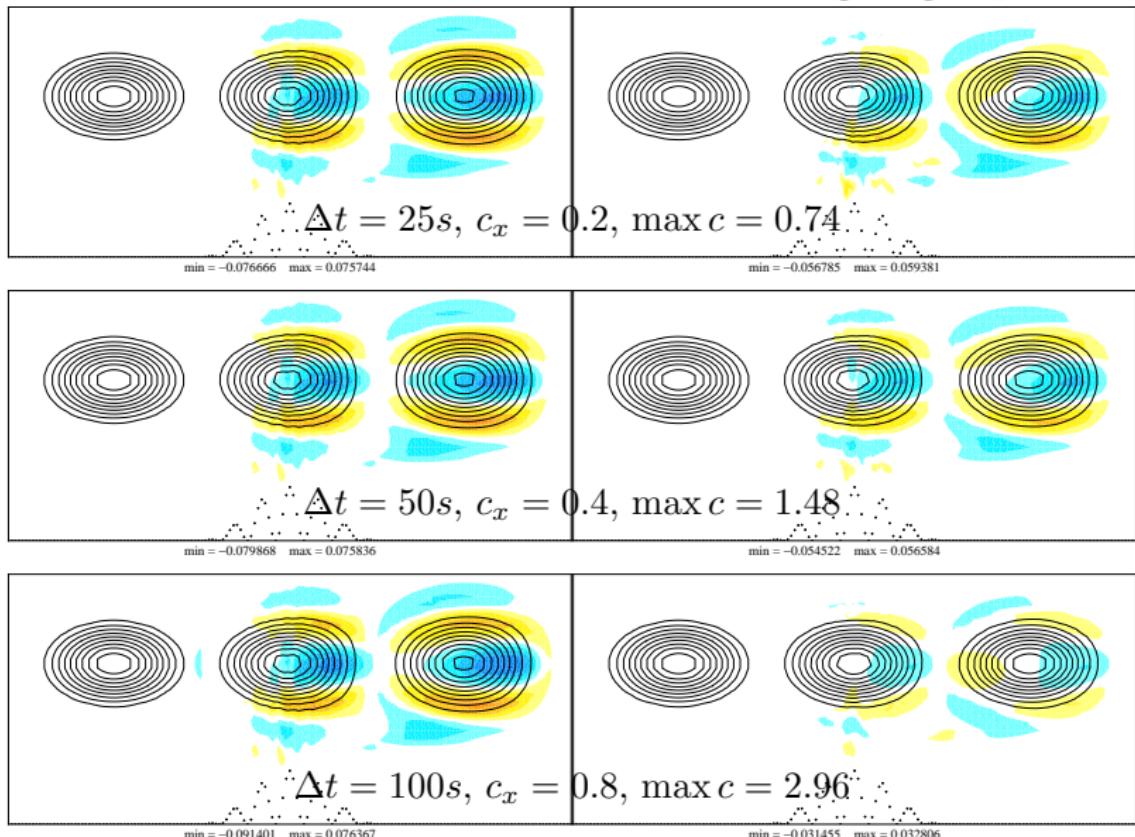
$\Delta t =$   
1000 s  
max  $c$ , 30  
mode  $c$ , 8  
 $c_d = 8.7$

unstable



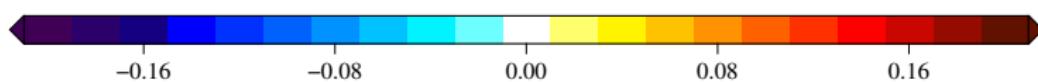
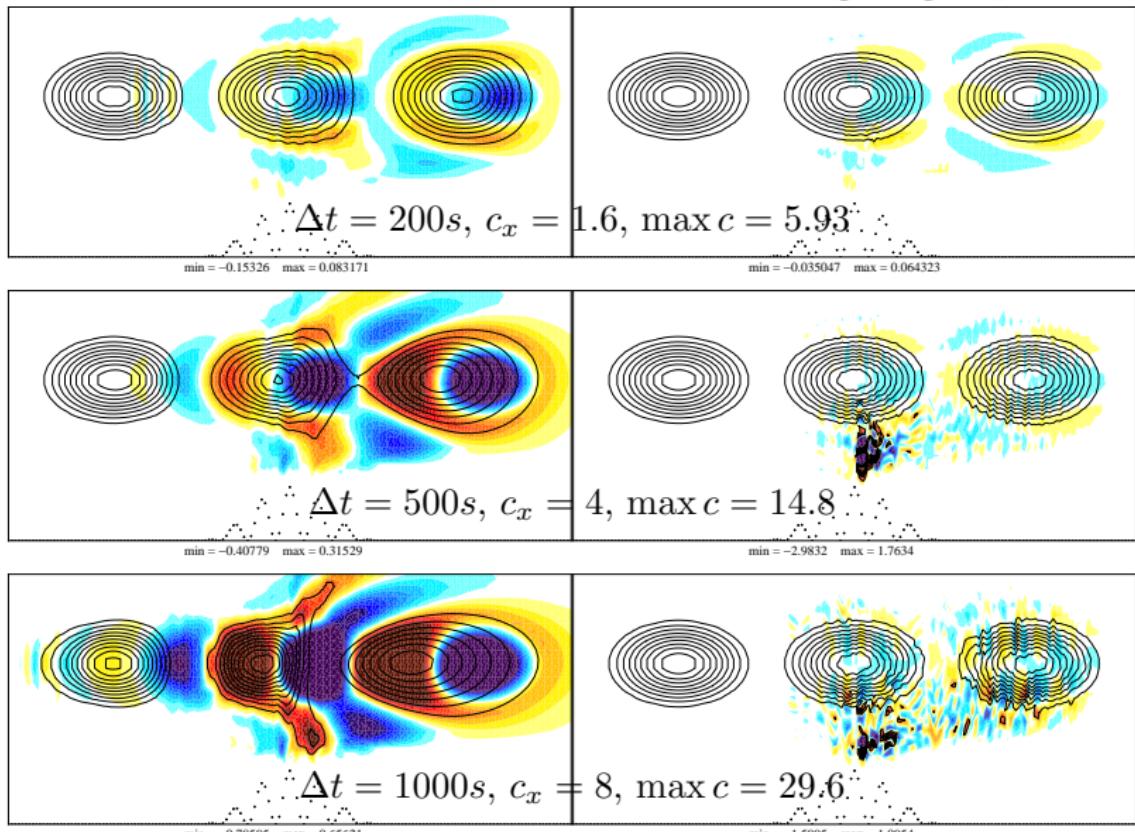
Contours from 0.1 to 0.9 every 0.1. Errors shaded.

# Multi-dimensional MOL Flux-form semi-Lagrangian, multi-d



Contours from 0.1 to 0.9 every 0.1. Errors shaded.

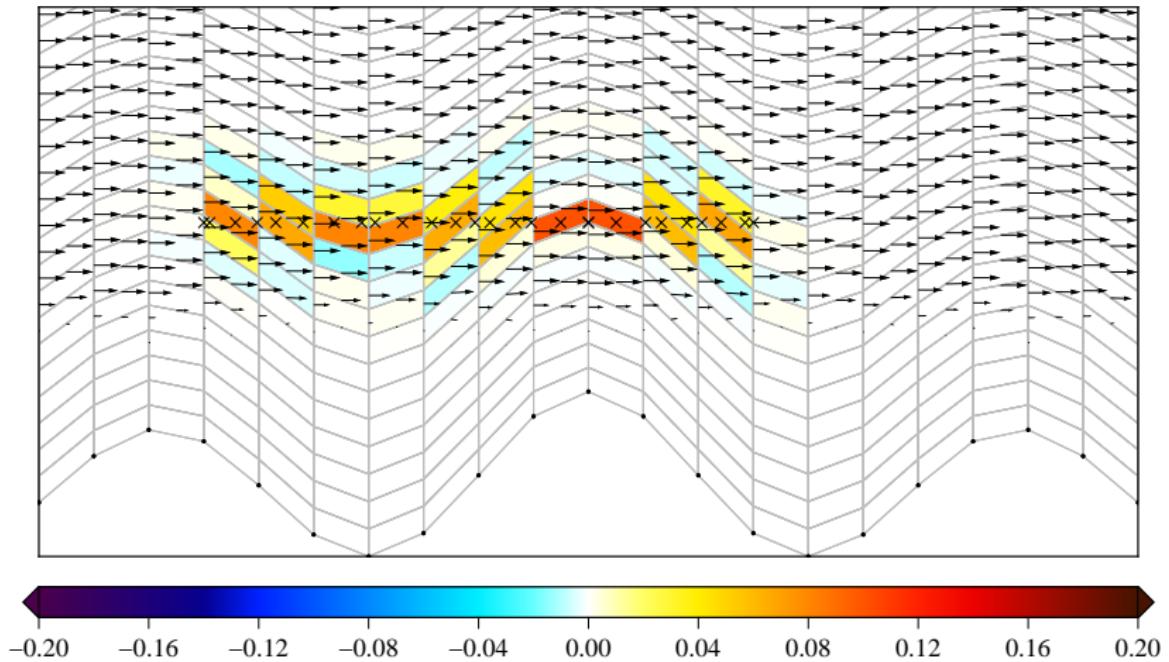
# Multi-dimensional MOL Flux-form semi-Lagrangian, multi-d



Contours from 0.1 to 0.9 every 0.1. Errors shaded.

# Flux Form semi-Lagrangian Stencil over Orography

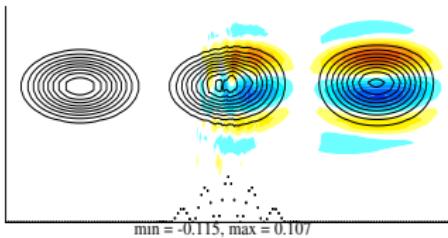
Points and weights for integrating along a trajectory.  $c_x = 8$ ,  $\max c = 29.6$



# Raise the mountain to $h_0 = 6\text{km}$

PPM, COSMIC splitting

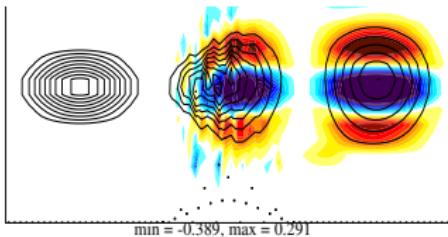
$\Delta t = 25\text{ s}$   
max  $c$ , 0.74  
mode  $c$ , 0.2



Multi-dimensional (MOL)

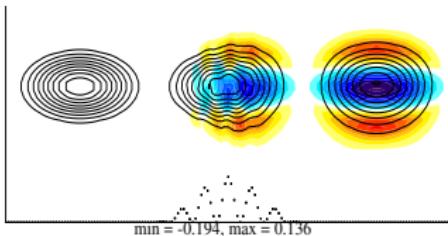
min = -0.076666, max = 0.075744

$\Delta t = 50\text{ s}$   
max  $c$ , 1.48  
mode  $c$ , 0.4

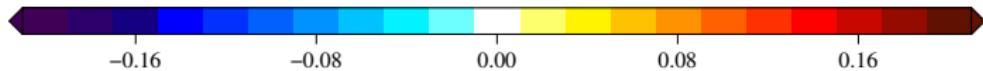


min = -0.079868, max = 0.075836

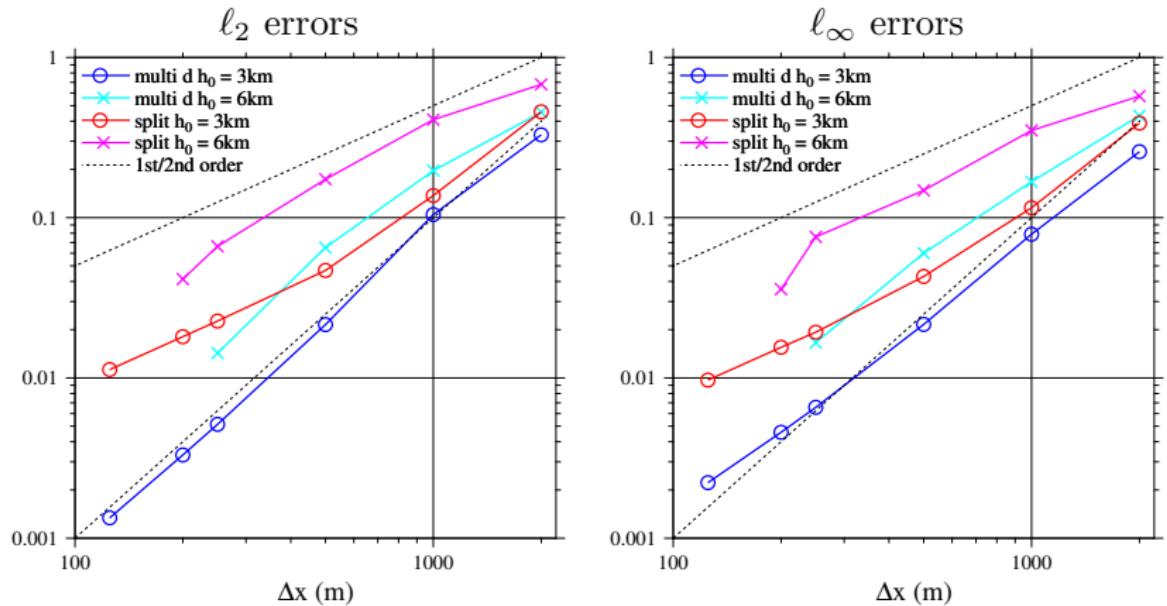
$\Delta t = 100\text{ s}$   
max  $c$ , 2.96  
mode  $c$ , 0.8



min = -0.091401, max = 0.076367

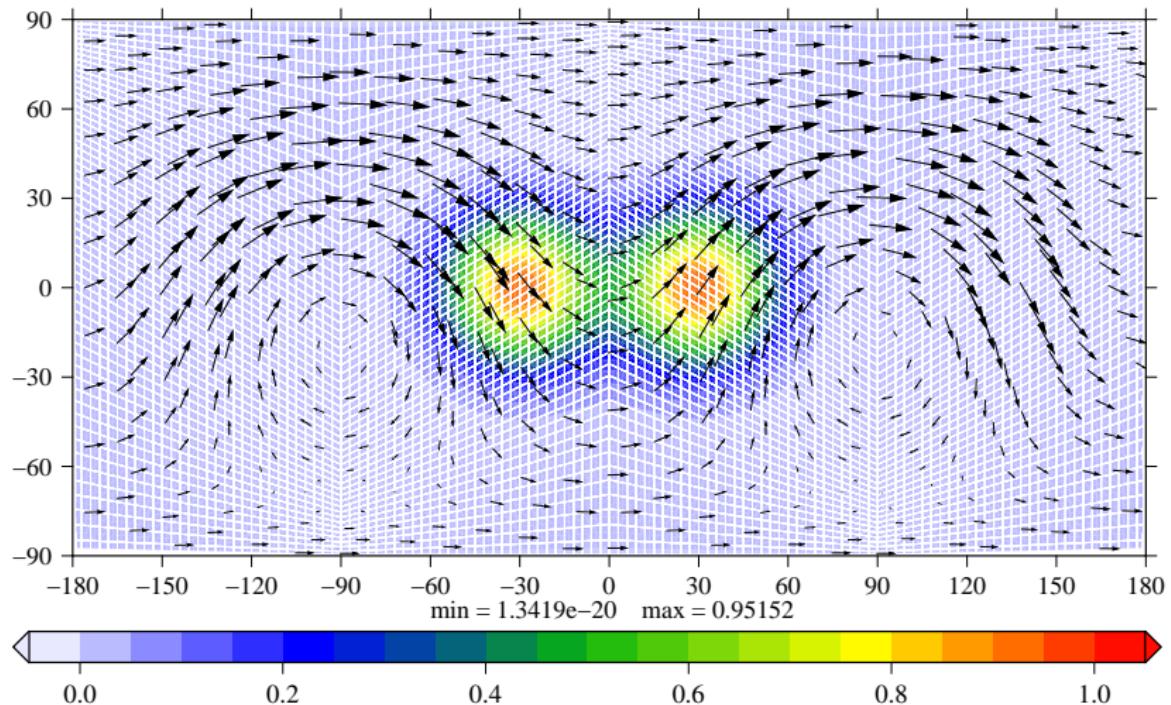


# Convergence with Resolution Over a Mountain



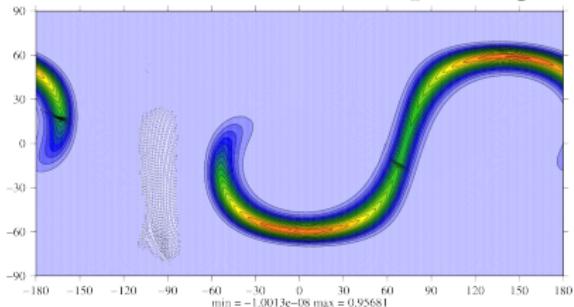
# Deformational Flow on a Plane

Using a non-orthogonal grid:

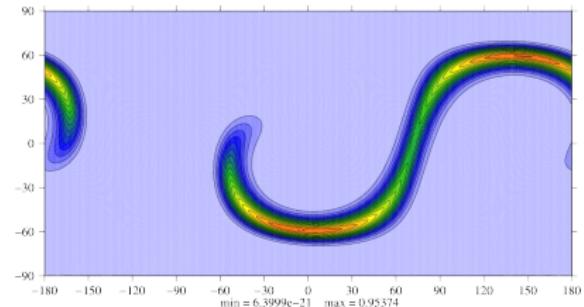


Results using  $480 \times 240$  cells,  $t = 1$  time unit

PPM with COSMIC splitting

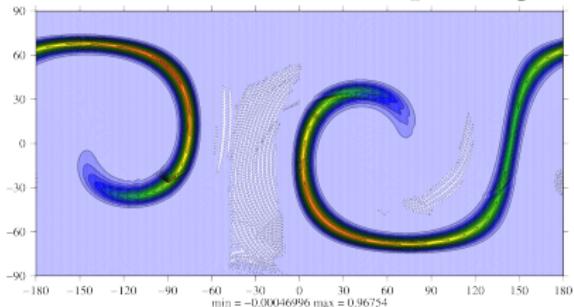


Multi-dimensional

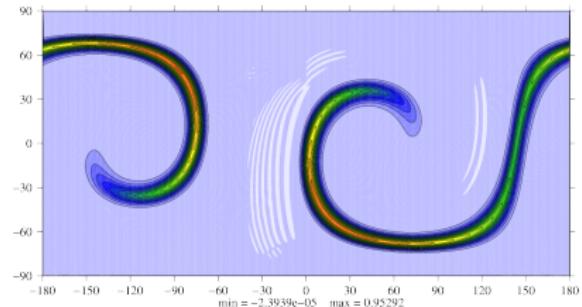


Results using  $480 \times 240$  cells,  $t = 2$  time unit

PPM with COSMIC splitting

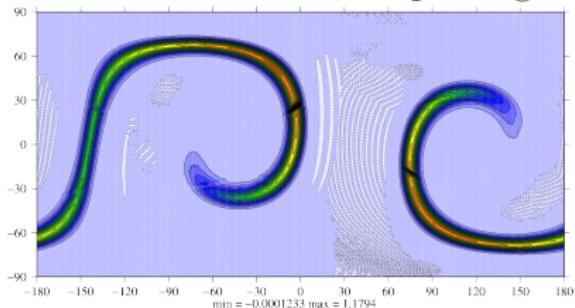


Multi-dimensional

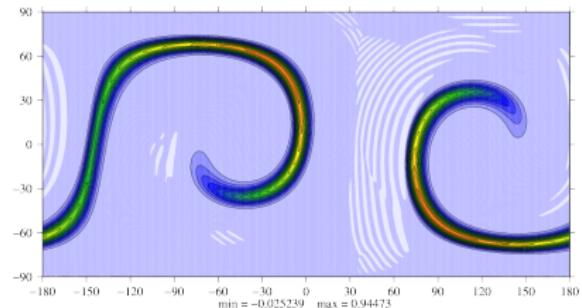


Results using  $480 \times 240$  cells,  $t = 3$  time unit

PPM with COSMIC splitting

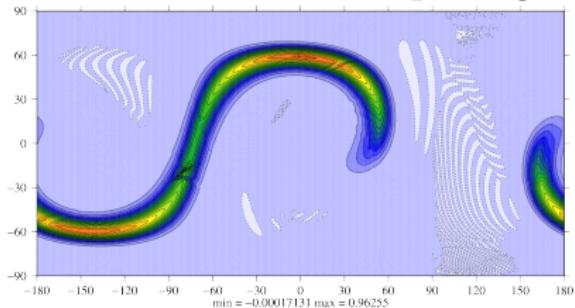


Multi-dimensional

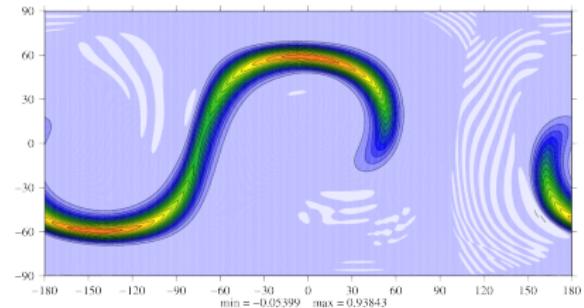


Results using  $480 \times 240$  cells,  $t = 4$  time unit

PPM with COSMIC splitting

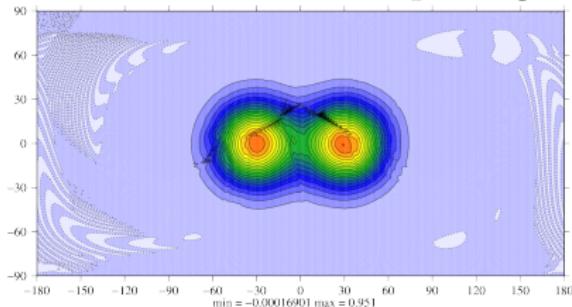


Multi-dimensional

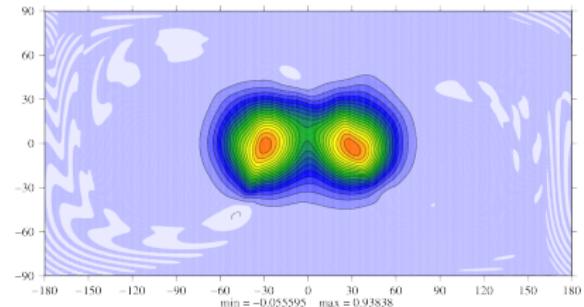


# Results using $480 \times 240$ cells, $t = 5$ time unit

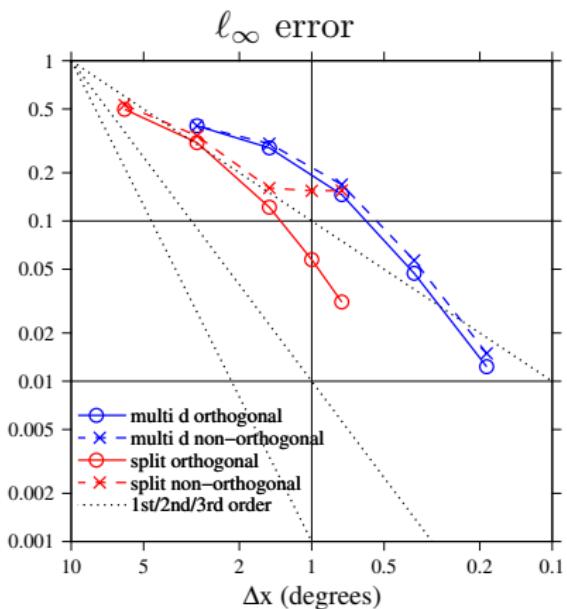
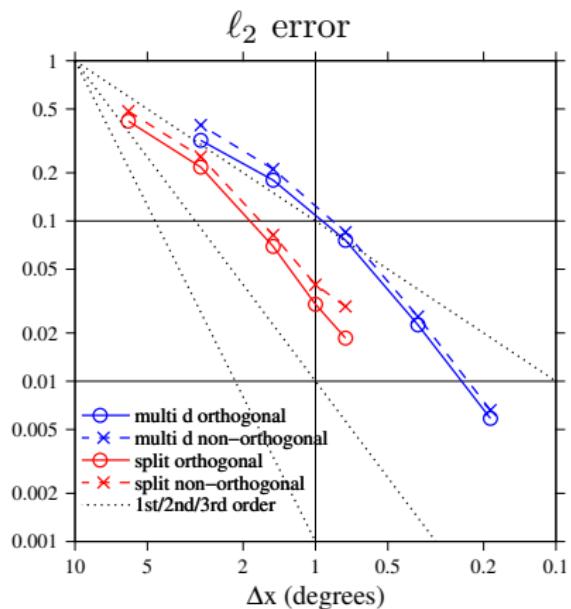
PPM with COSMIC splitting



Multi-dimensional



# Convergence with Resolution



# Conclusions

## Multi-dimensional flux form semi-Lagrangian

- ▶ Very expensive
  - ▶ Cost proportional to Courant number
  - ▶  $\text{int}(c)+1$ ,  $10 \times 13$  (in 2D) matrix inversions per face per time step for 2nd order accuracy with cubic interpolation
  - ▶ Computational geometry to find points in each cell along trajectory
- ▶ Stable for very long time-steps ( $c$  up to 30)
- ▶ Accurate for long time-step ( $c \leq 3$ )

# Conclusions

## Method of lines with Crank-Nicholson time-stepping

- ▶ Cost not strongly dependent on time-step
- ▶ Accurate on orthogonal and non-orthogonal grids
- ▶ Stable for very long time-steps ( $c$  up to 30) using deferred correction on 1st order upwind
- ▶ Phase lagging for  $c > 5$  but phase speed  $\not\rightarrow 0$  even for  $c$  up to 30

## PPM with COSMIC splitting

- ▶ More efficient and accurate on orthogonal grids
- ▶ Inaccuracies on non-orthogonal grids where non-orthogonality changes direction
  - ▶ excellent away from cubed-sphere edges
  - ▶ problems at cubed-sphere edges and over steep terrain (using terrain following grid)

- P. Colella and P. Woodward. The piecewise parabolic method (PPM) for gas-dynamical simulations. *J. Comput. Phys.*, 1984.
- B. Leonard, A. Lock, and M. MacVean. Conservative explicit unrestricted-time-step multidimensional constancy-preserving advection schemes. *Mon. Wea. Rev.*, 124(11):2585–2606, 1996.